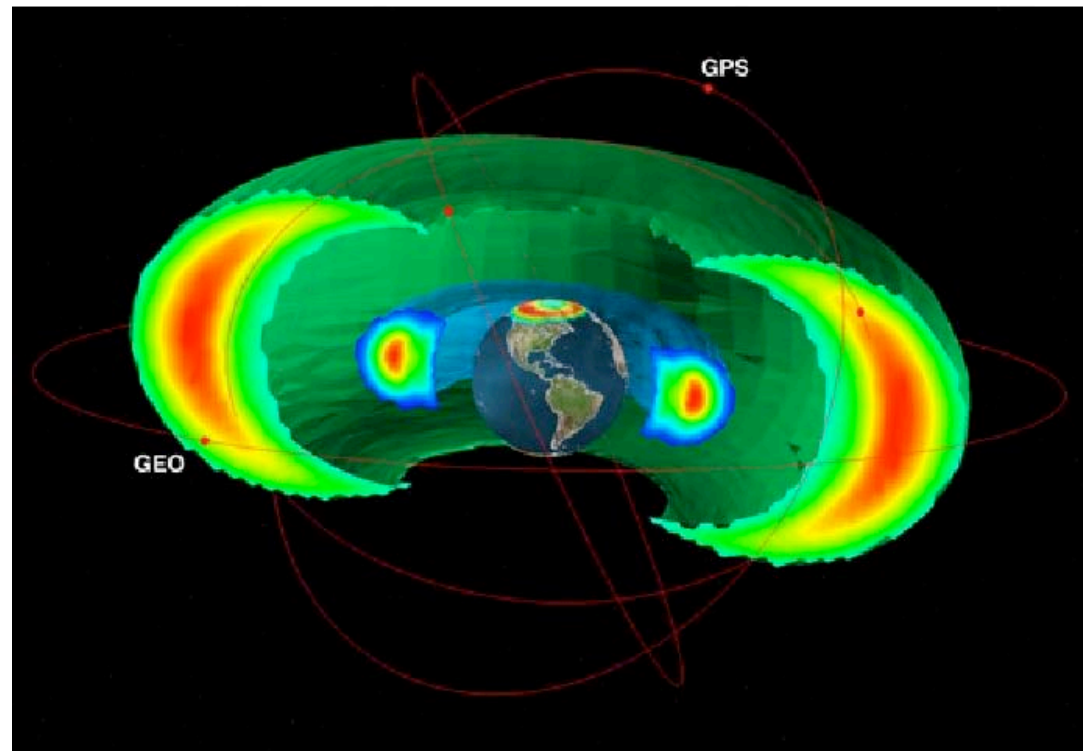
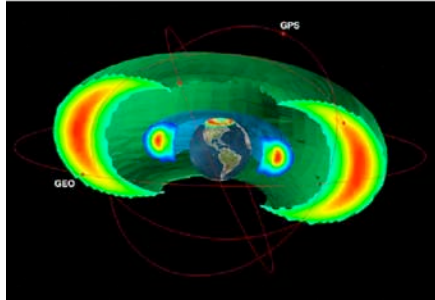


Introduction to the Radiation Belts



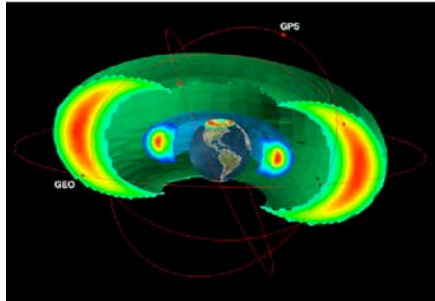
Wm. Robert Johnston

GEM Workshop - Midway, UT - 17 June 2007



Outline

- Introduction to radiation belts
- Trapped particle motions and adiabatic invariants
- Pitch angle
- Wave-particle interactions and the plasmasphere
- Source, loss, and diffusion mechanisms
- Observation platforms
- Conclusion



How far we've come...

- The magnetosphere in 1965:

ANTHROPOMORPHIC MAGNETOSPHERE

CHARACTERIZED BY

NOSE WHISTLERS

[VAN ALLEN'S] BELT

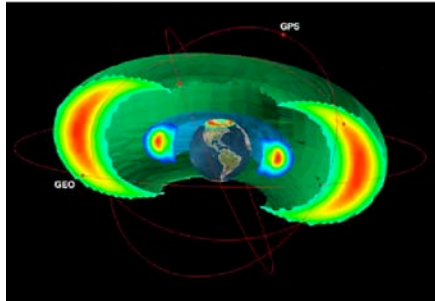
[CARPENTER'S] KNEE

AND

[NESS'] LONG TAIL

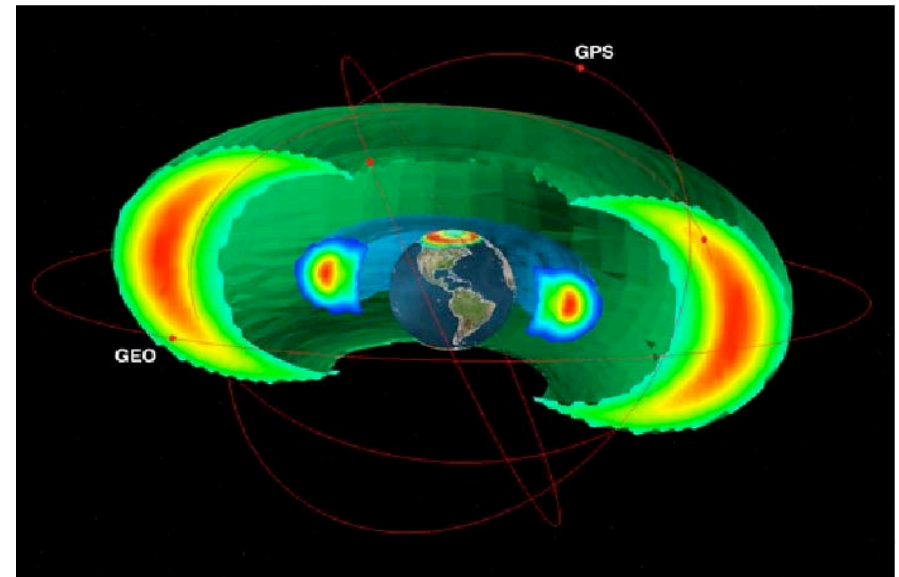


O'Brien, in Lemaire and Gringauz (1998)

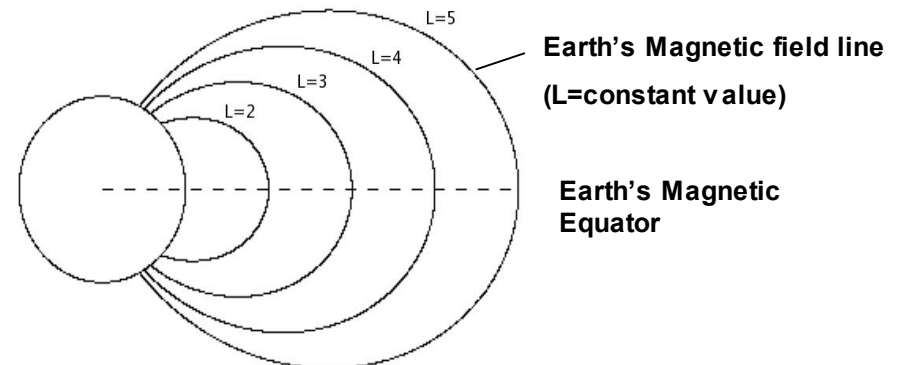


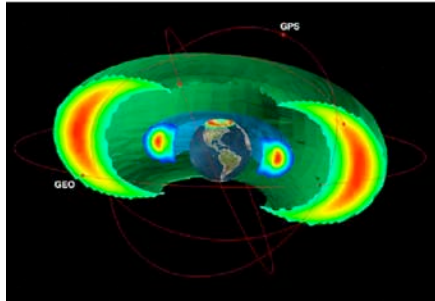
Radiation belt basics

- Radiation belts comprise energetic charged particles trapped by the Earth's magnetic field. (from keV to MeV)
- A given field line is described by its **L value** (*radial location, in R_E , of its intersection with magnetic equator*)
- Inner belt region:
 - Located at $L \sim 1.5-2$
 - Contains electrons, protons, and ions.
 - Very stable.
- Outer belt region:
 - Located at $L \sim 3-6$
 - Contains mostly electrons.
 - Very dynamic.
- Slot region: lower radiation region between the belts

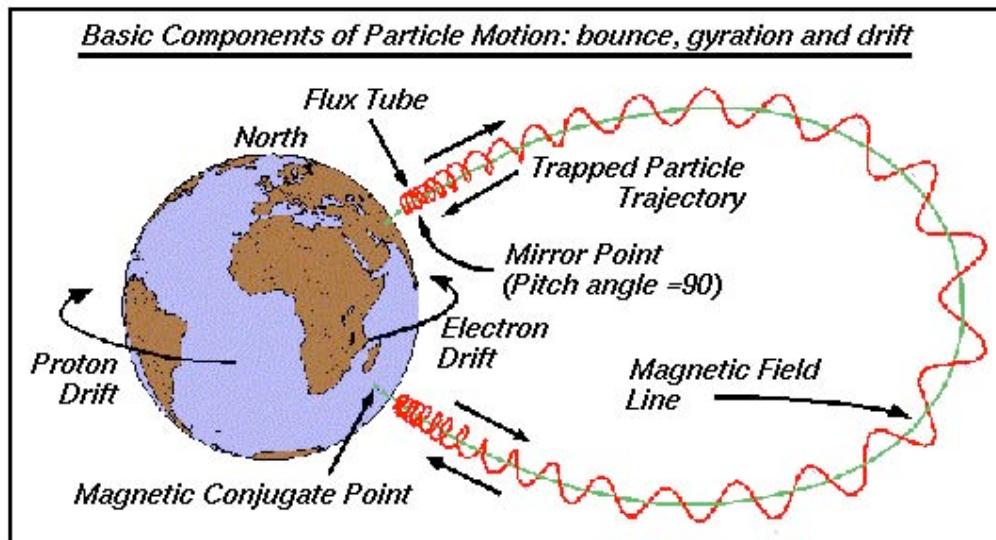


NASA





Periodic motions of trapped particles (1)



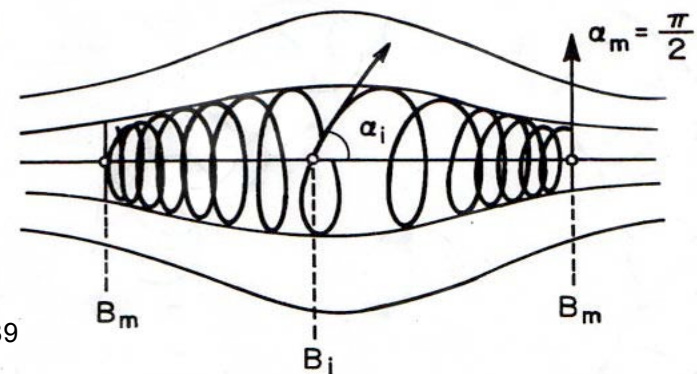
ESA

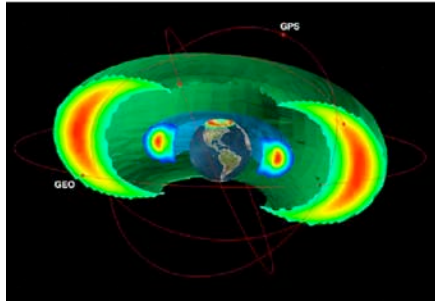
- Three types of periodic motion of trapped particles
 - gyro motion
 - bounce motion
 - drift motion
- Each motion has an associated adiabatic invariant

- Gyro motion:
 - $V \times B$ acceleration leads to gyro motion about field lines
 - frequencies \sim kHz
 - associated 1st invariant μ , relativistic magnetic moment:

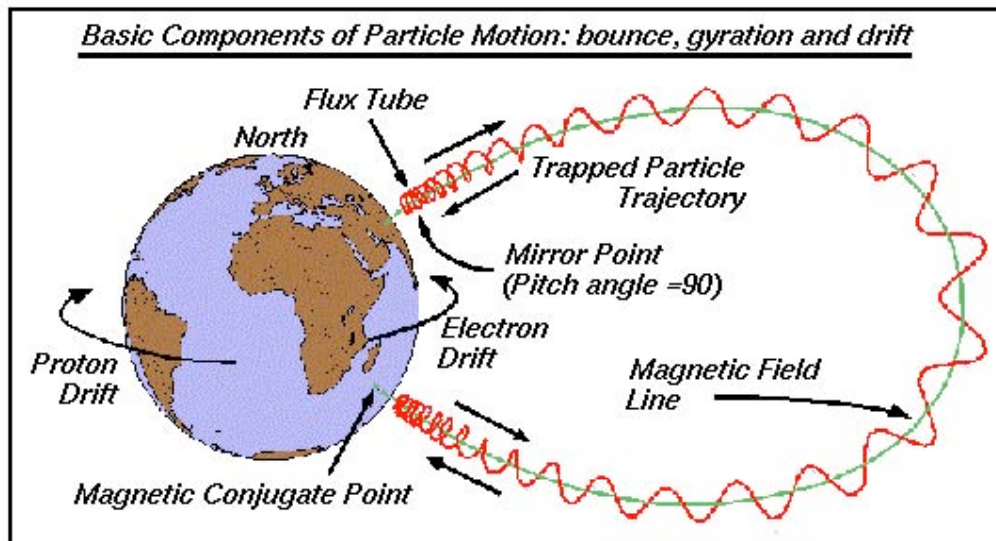
$$\mu = \frac{p^2 \sin^2 \alpha}{2m_0 B}$$

pitch angle α : $\tan \alpha = \frac{V_{\perp}}{V_{\parallel}}$



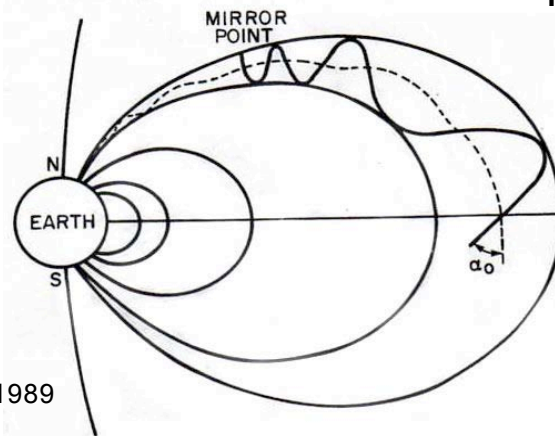


Periodic motions of trapped particles (2)



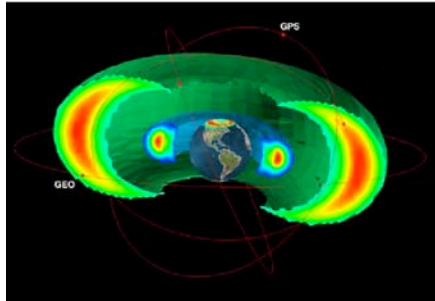
- Bounce motion:
 - As a particle gyrates down a field line, the pitch angle increases as B increases
 - Motion along field line reverses when pitch angle reaches 90° (mirror point)
 - period ~sec
 - associated 2nd invariant J, longitudinal invariant:

ESA

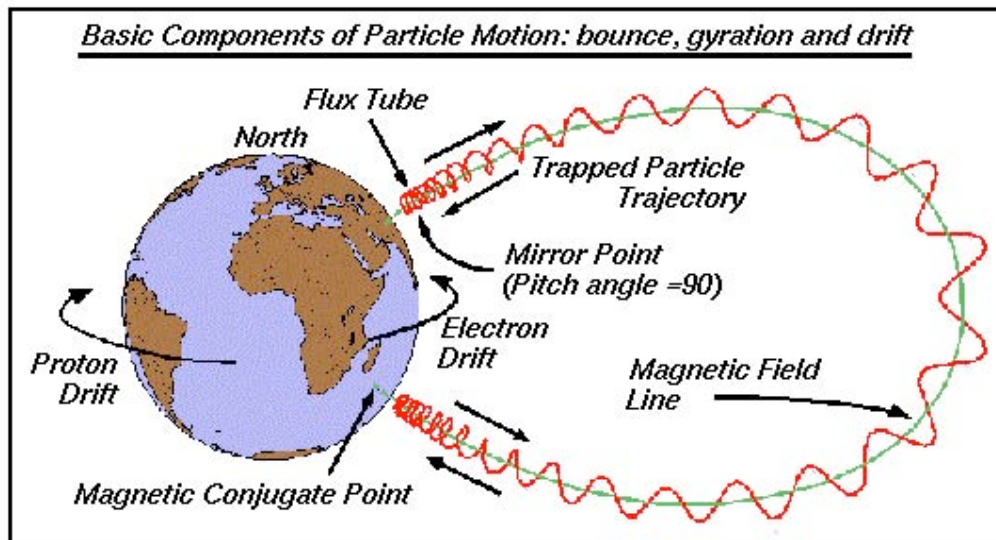


$$J = \int_{-l_m}^{+l_m} p_{\parallel} dl$$

Spjeldvik and Rothwell, 1989



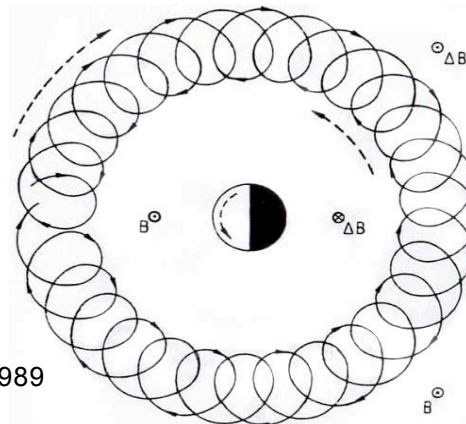
Periodic motions of trapped particles (3)



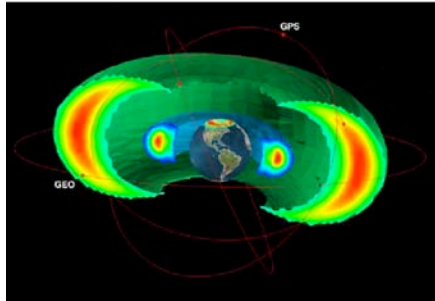
ESA

- Drift motion:
 - Gradient in magnetic field leads to drift motion around Earth: east for electrons, west for protons/ions
 - period ~minutes
 - associated 3rd invariant ϕ , magnetic flux:

$$\Phi = -\frac{2\pi B_E R_E^2}{L}$$

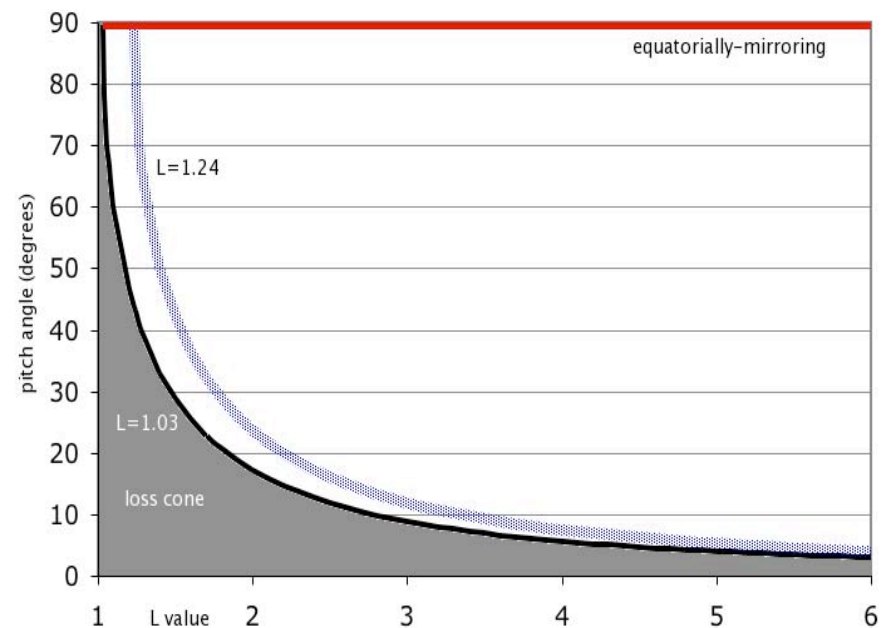
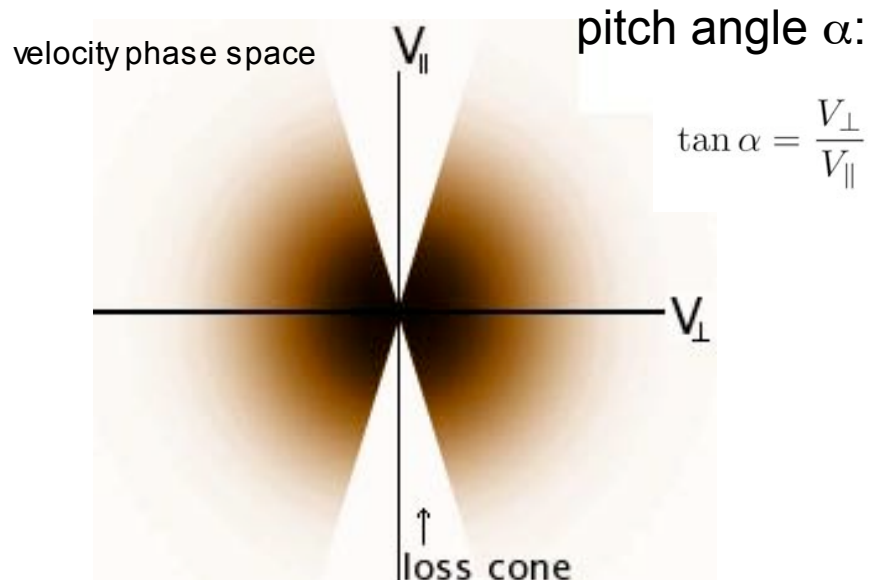


Spjeldvik and Rothwell, 1989

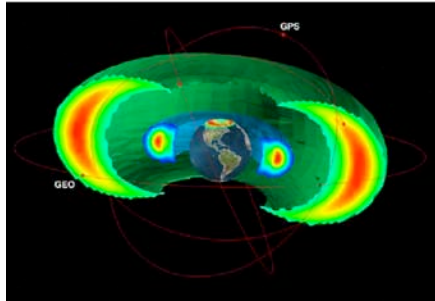


Pitch angle dependence

- Radiation belt populations are necessarily nonisotropic.
- Illustrated by nonisotropic distribution in velocity phase space:

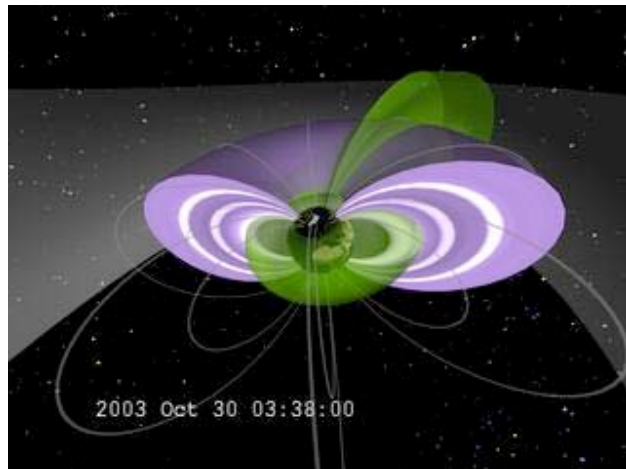


- Figure shows range of equatorial pitch angle values sustainable for mirroring particles.



Plasmasphere

- Plasmasphere--a torus of cold (~ 1 eV), dense (10 - 10^3 cm^{-3}) plasma trapped on field lines in corotation region of the inner magnetosphere
 - outer boundary (plasmopause) tends to correlate with inner boundary of outer radiation belt
 - typically extends to $L=3$ - 5 , but can be very structured and dynamic



NASA/GSFC

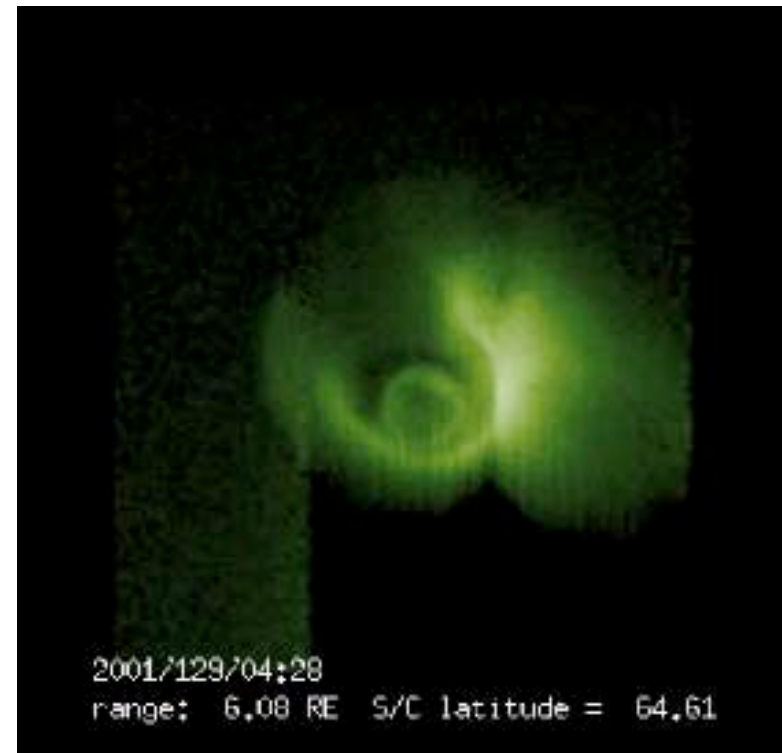
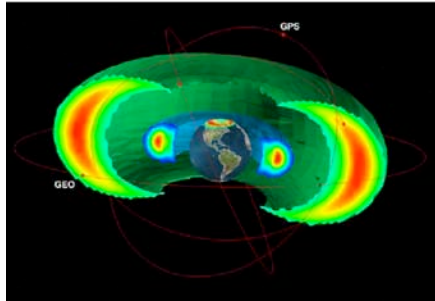
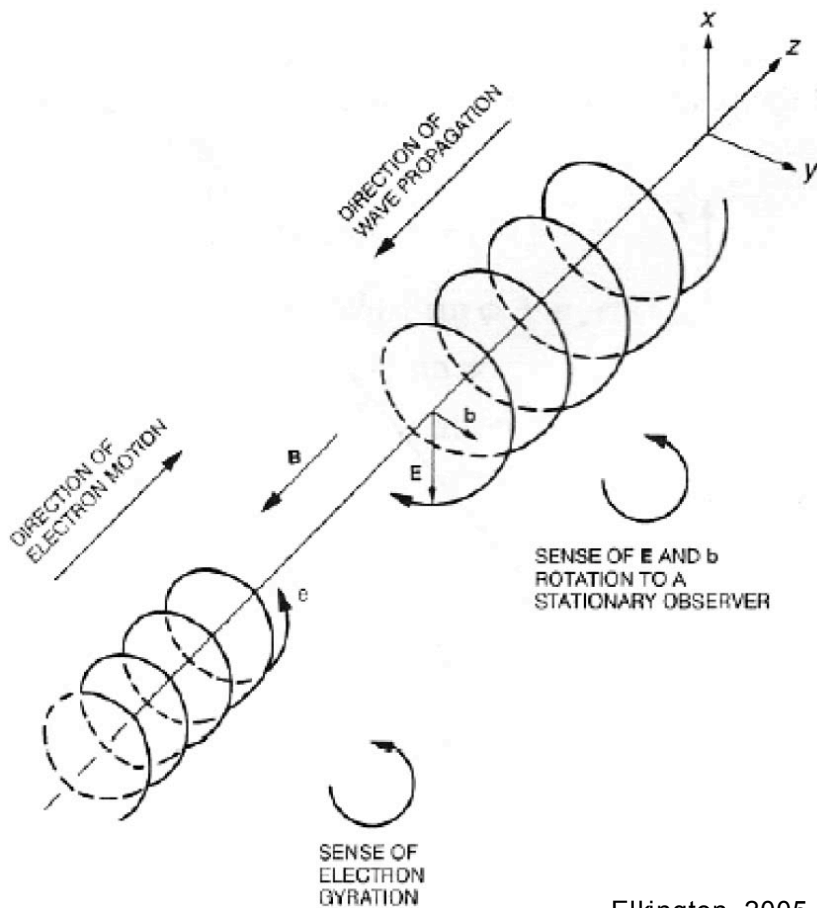


IMAGE EUV web site

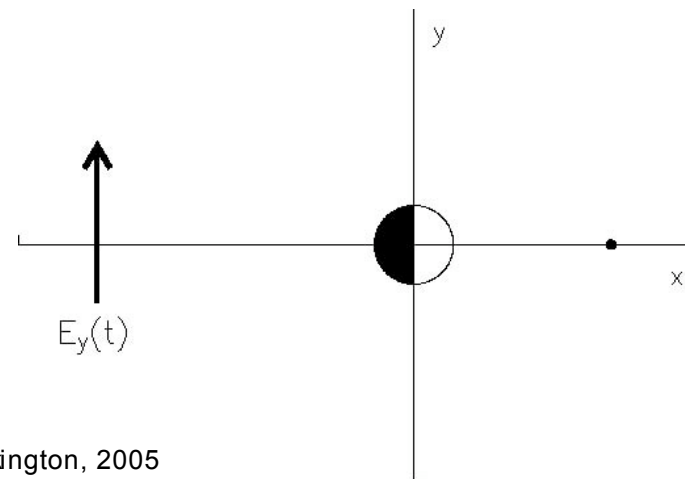


Wave-particle interactions

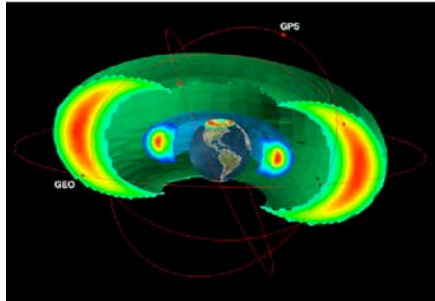


Elkington, 2005

- Wave-particle interactions: resonances between periodic particle motion and EM waves can energize or scatter particles
 - Whistler waves
 - ULF waves

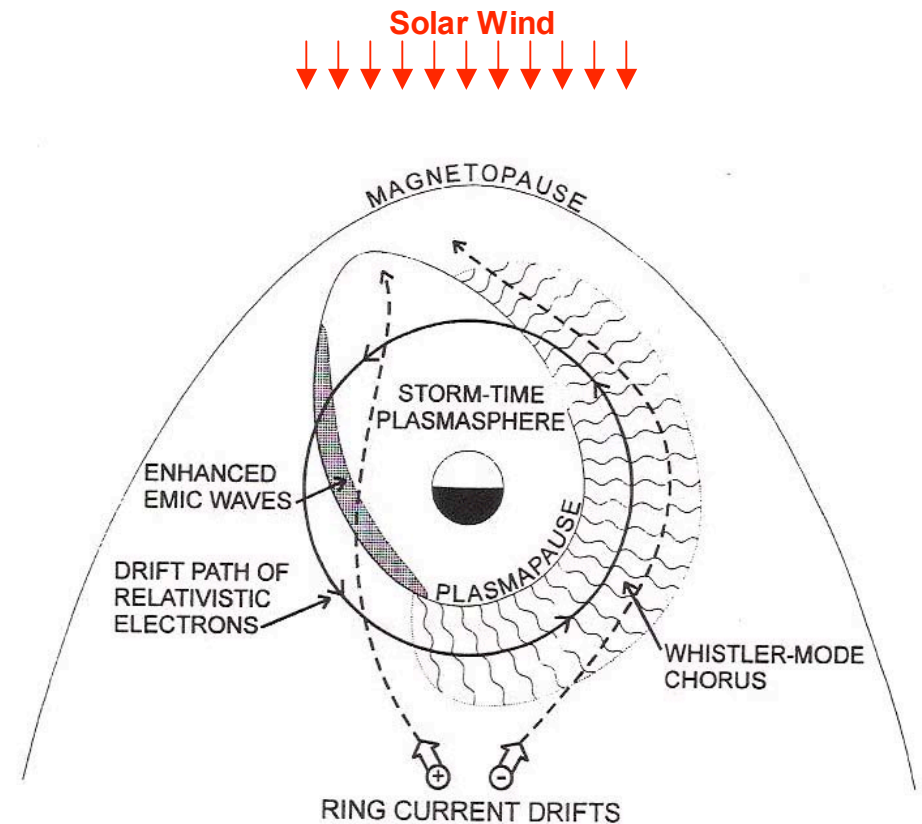
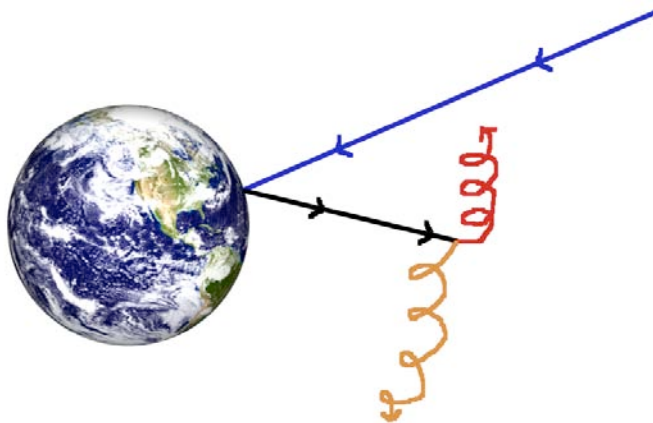


Elkington, 2005

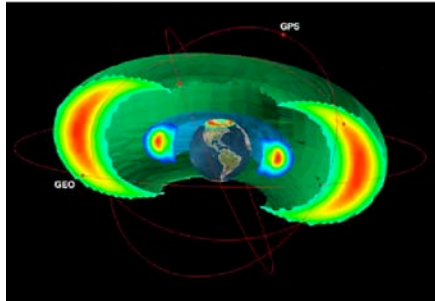


Sources and energization mechanisms

- Sources include solar wind via outer magnetosphere or from plasmashet plasma
- These particles energized by wave-particle interactions (e.g. whistler waves), crosstail E field fluctuations
- Cosmic ray albedo neutrons
 - cosmic rays --> n --> H⁺ and e⁻



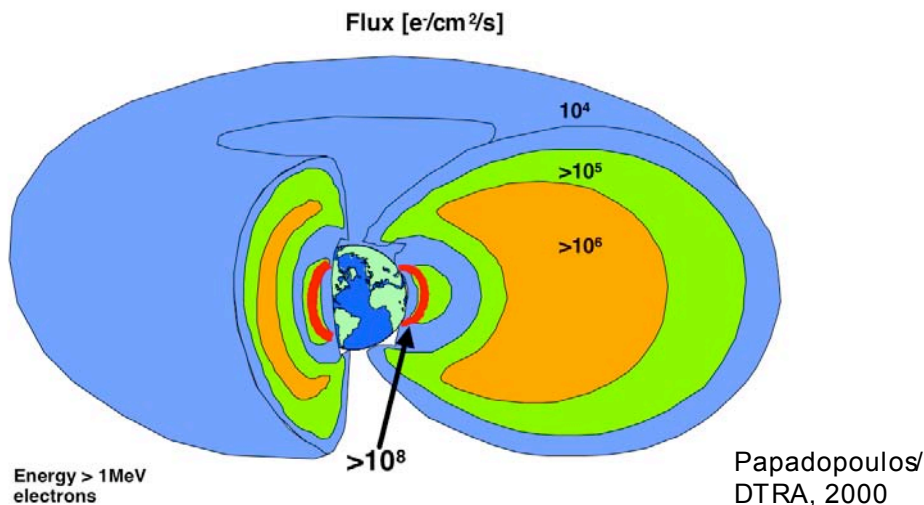
Summers et al., 1998



Man-made sources

- High altitude nuclear explosions can produce artificial radiation belts
 - several US, Soviet tests in 1958-1962 produced short-lived belts inside the inner belt

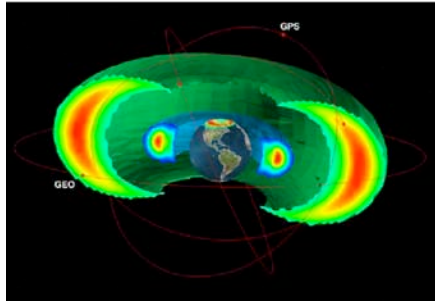
Natural and Enhanced Electron Population
One Day After Burst Over Korea



Starfish,
1962,
1.4 mt,
400 km alt.

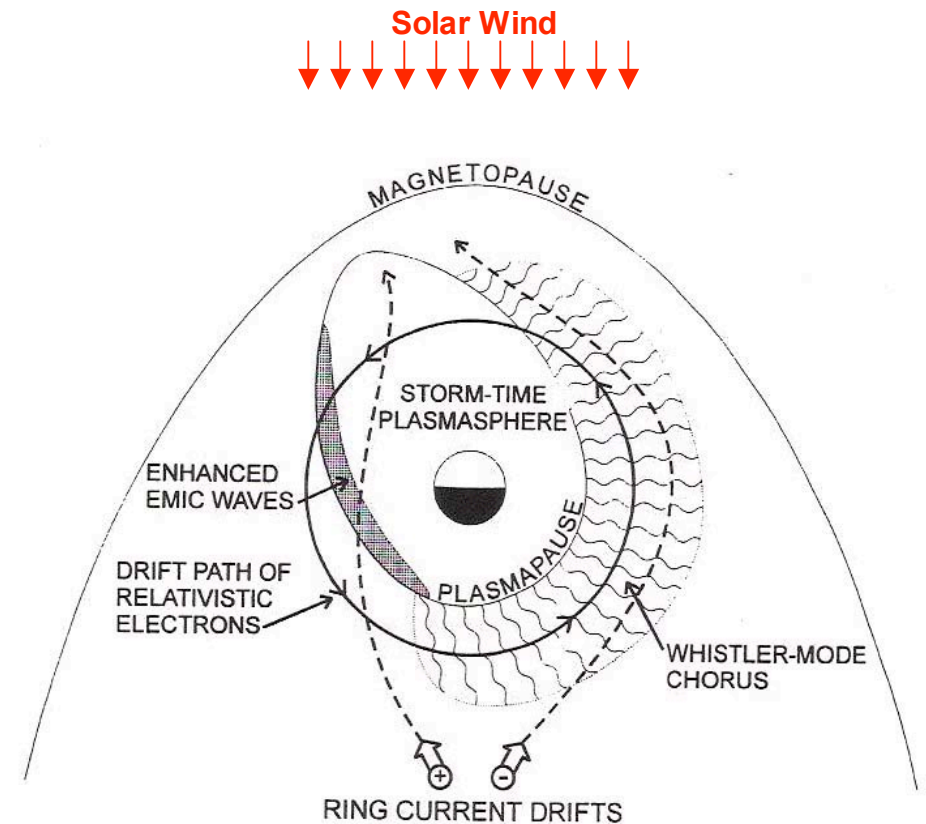
Nuclear Weapon
Archive, 2005

- Currently a national security concern given our dependence on space assets

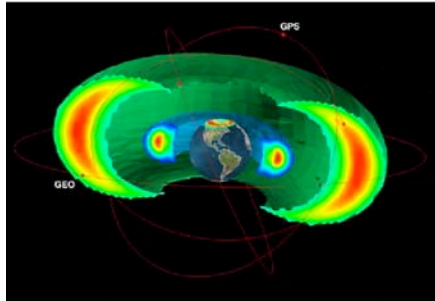


Loss mechanisms

- Anything that scatters particles into loss cone in phase space
 - such particles will collide with atmosphere
- Coulomb collisions with cold charged particles in plasmasphere, ionosphere
- Enhanced EMIC waves inside plasmapause
- Magnetopause shadowing
 - loss of particles with orbits carrying them outside the magnetopause

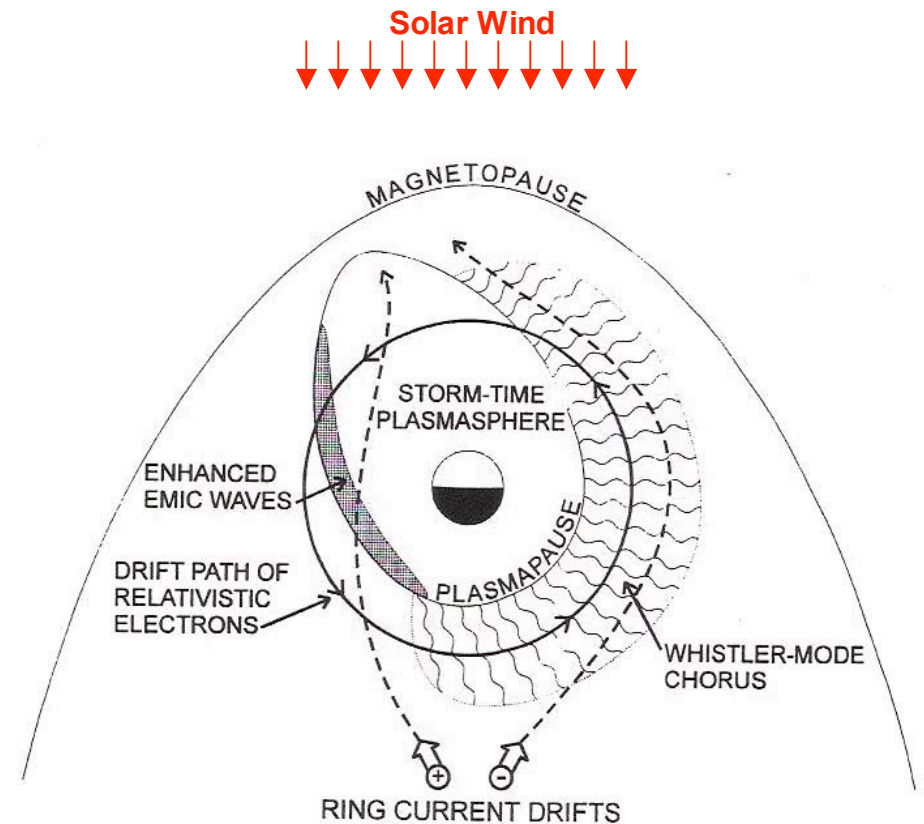


Summers et al., 1998

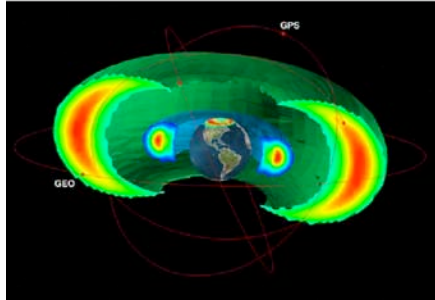


Diffusion mechanisms

- Wave-particle interactions
 - whistler chorus
 - EMIC waves
- Fluctuations in magnetospheric electric field



Summers et al., 1998



Diffusion equations and phase spaces

- Evolution of particle population described by **diffusion equation**:

$$\text{rate of change in flux} = \text{sources} - \text{losses} + \text{diffusion terms}$$

- What phase space to use to model evolution?

basic position-
momentum space:

x, y, z, v_x, v_y, v_z

(hard to use)

adiabatic
invariants:

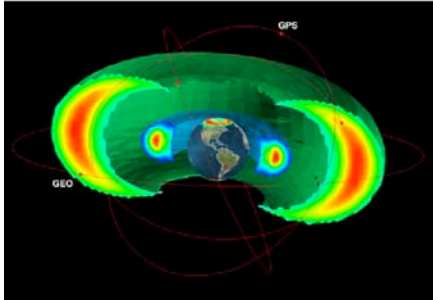
μ, J, ϕ

(easy to use)

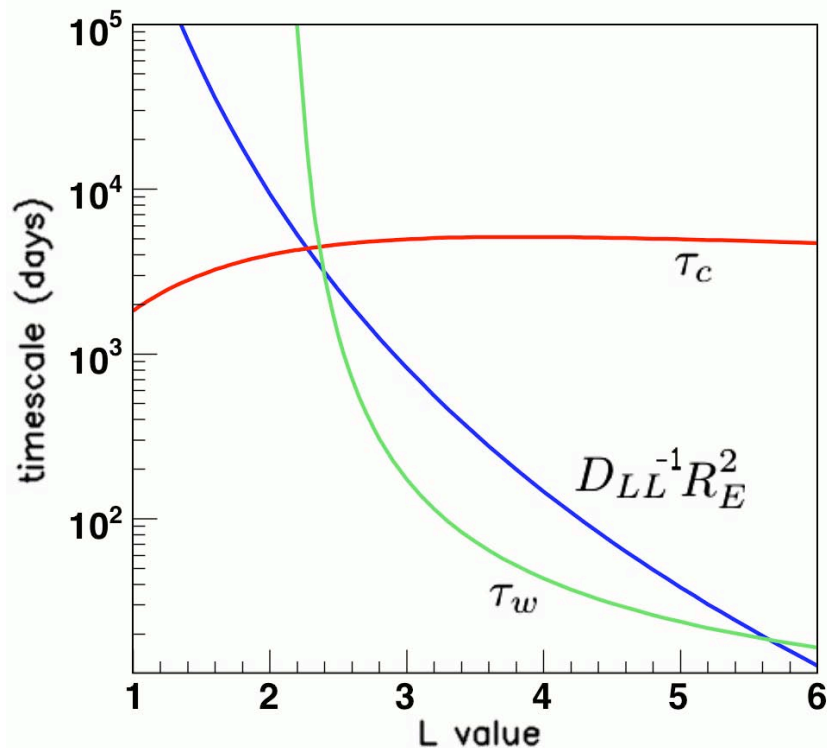
observables:

ϵ, α_0, L

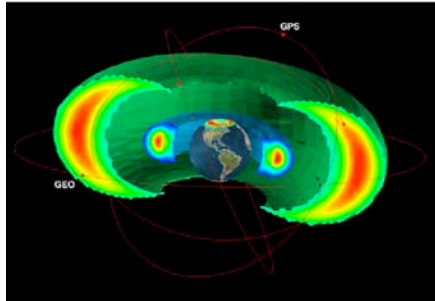
(for interpretation)



Why there are two electron belts

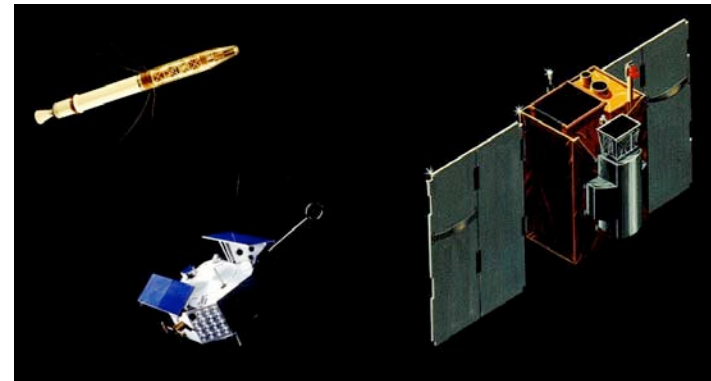


- plot shows timescales for fixed $\mu=30$ MeV/G (after Lyons and Thorne, 1973)
- D_{LL} drives inward diffusion, faster at large L
- whistler losses faster than replacement by diffusion in slot region
- those particles that reach low L have lifetimes of years

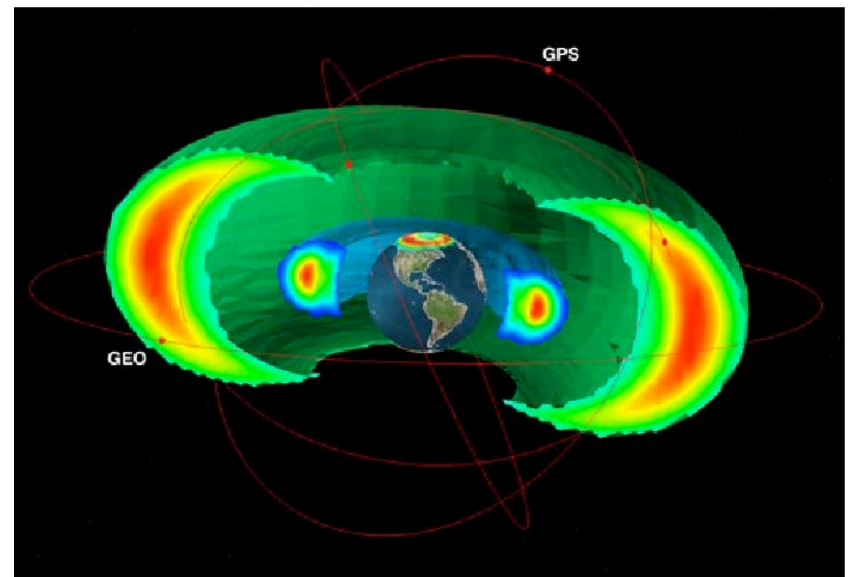


Illustrative satellites

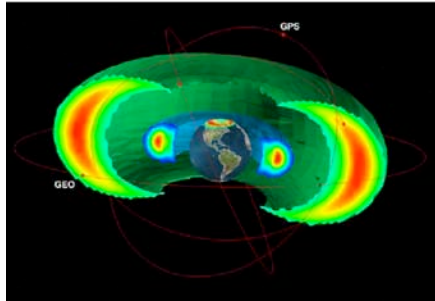
- Explorer 1/3 (1958)
 - low Earth orbit, eccentric
 - geiger counter
- later satellites: multiple particle detectors, pitch angle info if spinning
- GOES (multiple, 1975-now)
 - geosynchronous orbit
- CRRES (1990-91)
 - eccentric orbit
- SAMPEX (1991-now)
 - low Earth orbit



NASA

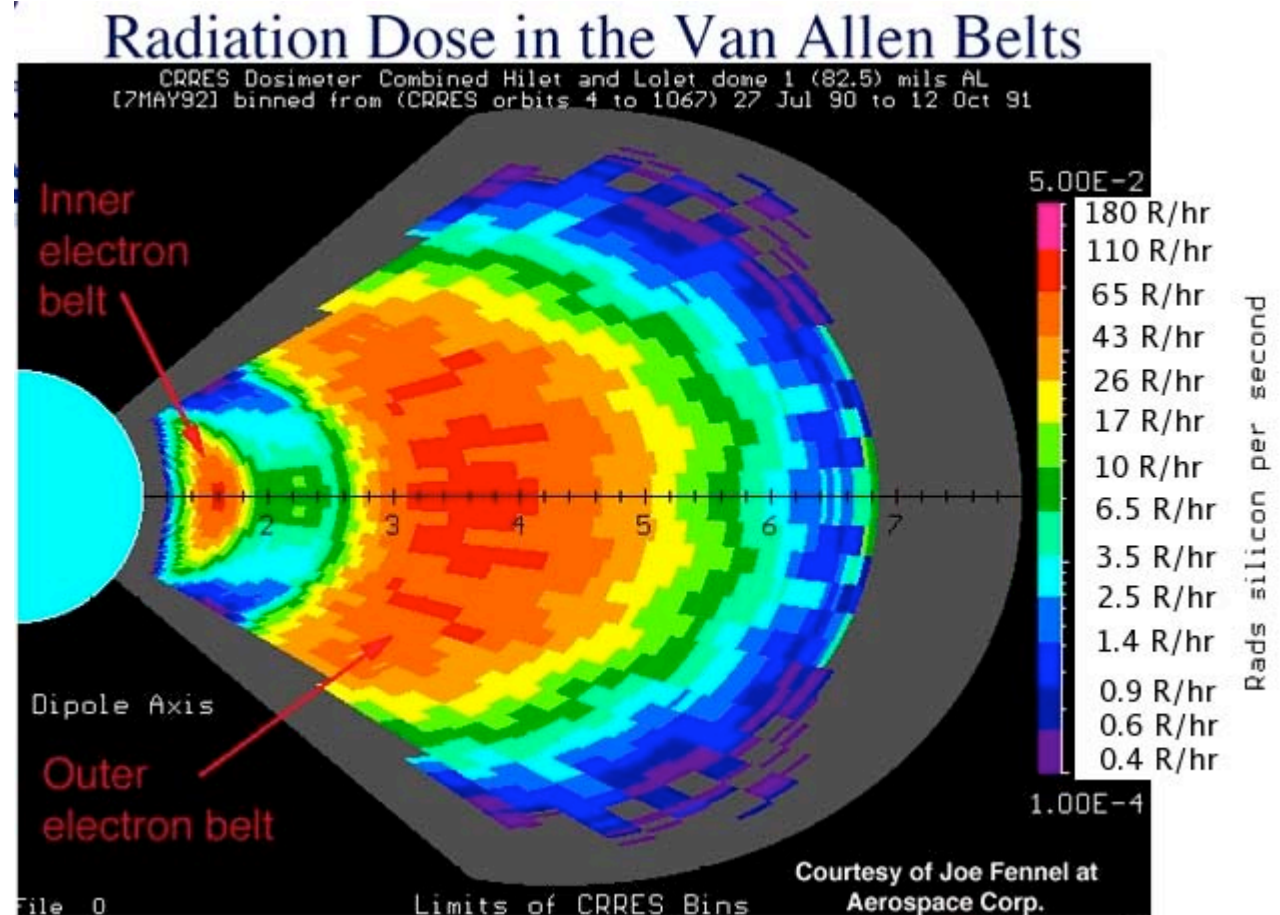


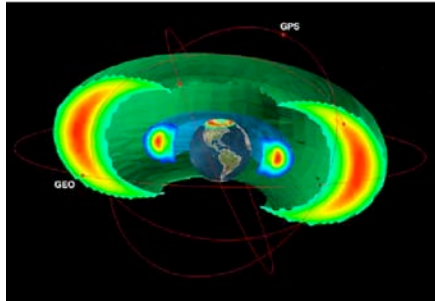
NASA



Radiation fluxes from CRRES

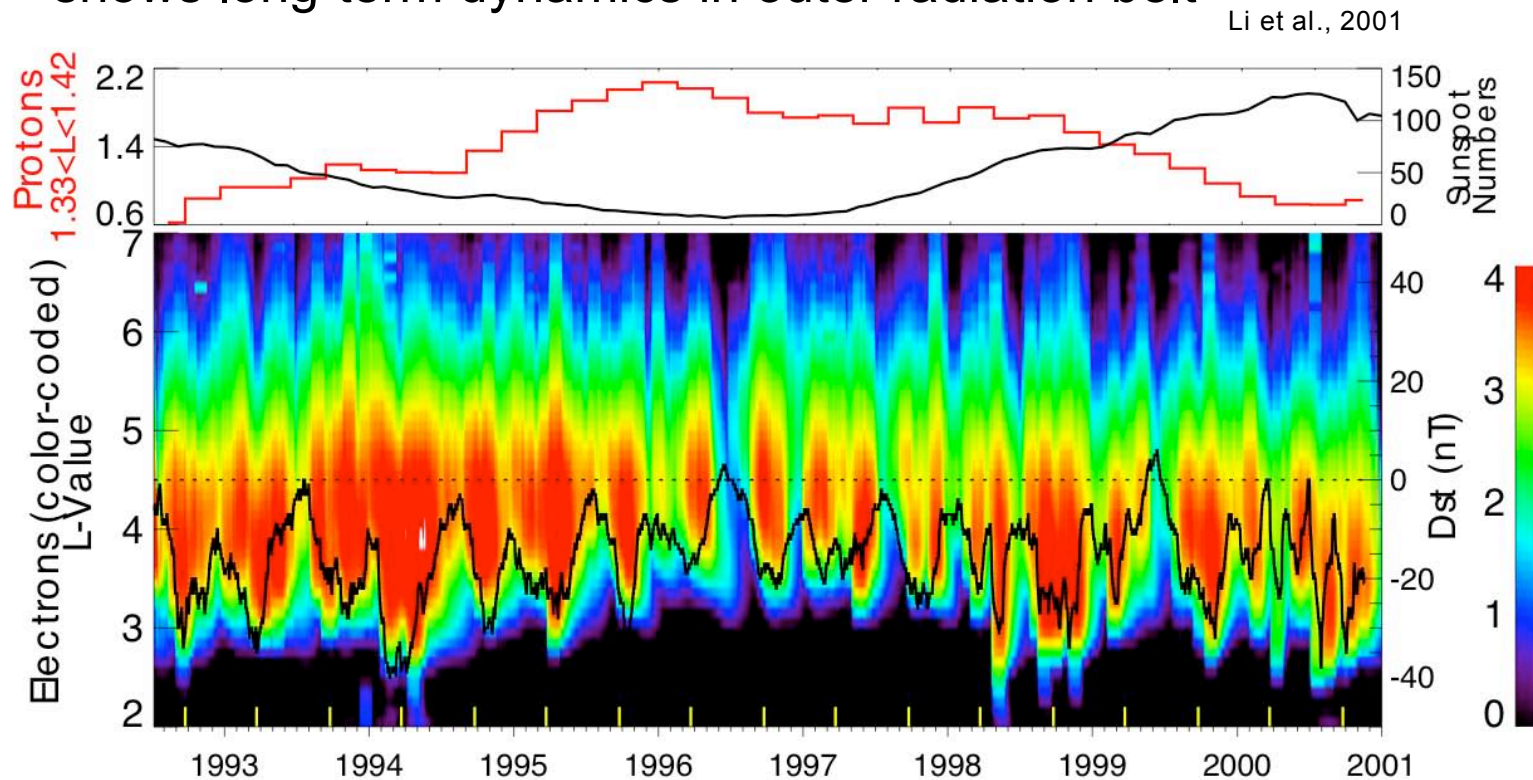
- CRRES= Combined Release and Radiation Effects Satellite
- radiation flux observations from CRRES, 1990-91
- scale converted to rads/hour

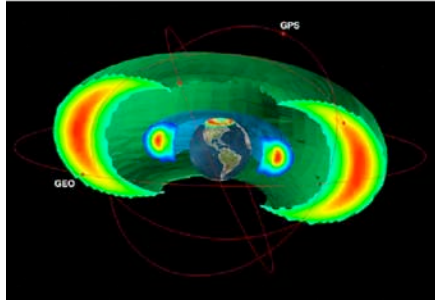




Long term dynamics from SAMPEX

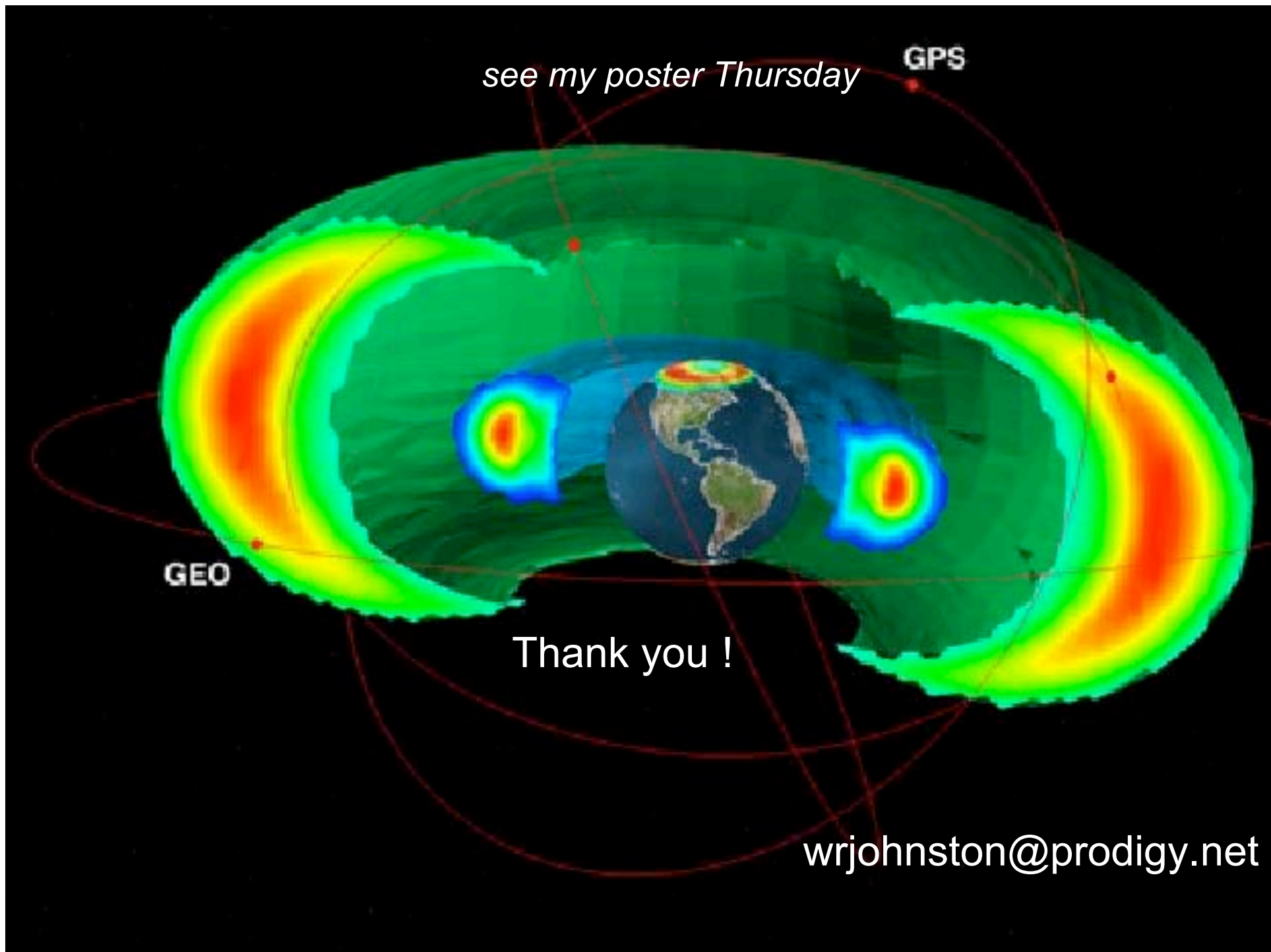
- SAMPEX=Solar Anomalous and Magnetospheric Particle Explorer
- SAMPEX observations over most of a solar cycle
- shows long-term dynamics in outer radiation belt





Conclusion

- Study of radiation belts is a rich topic with connections to many space physics subfields.
- Understanding of radiation belts is important to space operations, both manned and unmanned.
- Currently a “hot” topic from many different perspectives!



see my poster Thursday

GPS

GEO

Thank you !

wrjohnston@prodigy.net